# Dark Matter, Dark Energy, & The Fate of the Universe

#### Chapter 16

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### Vark Matter and Vark Energy?

\* This is our current situation

 Gravity (from luminous & dark matter) holds galaxies together and attracts each galaxy to all others and fights against expansion

2. Yet space is expanding and carries galaxies away from one another, weakening gravitational forces

### Dark Matter and Dark Energy?...

#### \* The crucial question

# \* how will that affect the future of our Universe?

\* Are there variations to that future, depending on the amount of gravity versus dark energy?

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### Unseen Influences in the Cosmos

\* Careful studies have shown that

a. galaxies do not have enough mass (that we can see) to hold together (yet they do...) hence more matter is needed

 b. the expansion of the Universe is accelerating hence some energy is doing this



 Galaxies are not flying apart as they rotate: there has to be more mass in and around each galaxy and we cannot see it

# An Additional Energy?

- \* When Einstein wrote the General Theory of Relativity he believed the Universe was static
- But his equations showed that the Universe was either expanding or contracting (with a constant rate)
- ★ He then introduced a quantity that fought against gravity: the cosmological constant, A

### General Relativity

These terms (R and g) describe the curvature of spacetime (they tell matter and energy how to move)

This term (T, the stress-energy tensor) describes the distribution of matter and energy (it tells spacetime how to curve)

 $\mathbf{R}_{\mu\nu} - \frac{1}{2} \mathbf{g}_{\mu\nu} \mathbf{R} + \mathbf{g}_{\mu\nu} \boldsymbol{\Lambda}$ 



The cosmological constant  $(\Lambda)$  behaves gravitationally like matter and energy except that it has negative pressure.

### Einstein's Cosmological Constant

\* When Hubble showed the Universe was not static but expanding, Einstein removed the term out of the equations thinking he had made a blunder by forcing the Universe to be static

\* However, we now are capturing data showing that the Universe expansion rate is accelerating and the constant is back: empty space has an intrinsic energy

### Unseen Influences

- Park Matter: the invisible mass: a recently detected form of matter which does not interact with the electromagnetic force but whose presence is inferred from its gravitational influence
- \* Park Energy: the invisible energy: an unknown form of energy that seems to be the source of a repulsive force causing the expansion of the Universe to accelerate



# Contents of the Universe

~4.9%
~0.4%
~4.5%
~26.8%
~68.3%

### Snapshot

- Park matter is the name given to the invisible mass whose gravity governs the observed motions of stars and gas clouds
- Park energy has never been directly observed, but it has been proposed to exist because it seems the simplest way to explain a set of observed motions in the Universe
- \* Park energy is the name given to whatever may be causing the expansion of the Universe to accelerate

# What is the evidence for dark matter in galaxies?

# \* Let's compare how different systems rotate

- \* the Solar System
- \* a merry-go-round
- \* a spiral galaxy



**Rotation curve** 

A plot of orbital velocity versus orbital radius



Solar System's rotation curve declines because Sun has almost all the mass





Rotation curve

Rotation curve of Milky Way stays flat with distance

Mass must be more spread out than in Solar System

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#### Milky Way rotation curve Its flatness indicates that our galaxy's mass extends well past the Sun's orbit



#### Most of the Milky Way's mass seems to be dark matter!



The visible portion of a galaxy lies deep in the heart of a large halo of dark matter

The radius of this dark matter halo may be 10 times as large as the galaxy's halo of stars

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#### What about other galaxies?



We can measure rotation curves of other spiral galaxies using the Doppler shift of the 21-cm line of atomic hydrogen

Friday, April 26, 13

### Spiral galaxies all tend to have flat rotation curves indicating large amounts of dark matter



### Elliptical galaxies do not have enough atomic gas to measure their 21-cm radiation



Instead we measure the motion of their inner stars

Broadening of spectral lines in elliptical galaxies tells us how fast the stars are orbiting

These galaxies also have dark matter



- \* What would you conclude about a galaxy whose rotational velocity rises steadily with distance beyond the visible part of its disk?
  - A. Its mass is concentrated at the center
  - B. It rotates like the solar system
  - C. It's especially rich in dark matter
  - D. It's just like the Milky Way

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### What is the evidence for dark matter in clusters of galaxies?

matter	visible	dark	factor
within galaxy	10%	90%	<b>10x more</b> (10+90)/10
within galaxy cluster	27,	98%	<b>50x more</b> (2+98)/2

# Fritz Zwicky

\* He was one of the first astronomers to think of galaxy clusters as huge swarms of galaxies bound together by gravity





An eccentric personality, and strange ideas for the 1930s that proved to be correct decades later

#### Method 1: Orbits of galaxies in clusters



We can measure the velocities of galaxies in a cluster from their Doppler shifts

The mass we find from galaxy motions in a cluster is about 50 times larger than the mass in stars! (which means 2% visible, 98% invisible)

#### Method 2: Hot gas in clusters



X-ray observations detect lots of hot gas filling space between galaxies

Temperature of this hot gas (particle motions) depends on the total (visible and invisible) mass of the cluster:

85% dark matter 13% hot gas 2% stars

#### Method 3: Gravitational lensing



#### Gravitational lensing, the bending of light rays by gravity, can also tell us a cluster's mass

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#### A gravitational lens distorts our view of things behind it



#### A gravitational lens distorts our view of things behind it



### Direct Evidence of Dark Matter Found

- \* As recently as 2006, dark matter has gone from a theoretical matter to a real one:
- By looking at distortions due to gravitational lensing, the influence of dark matter has been directly detected in a collision of two galactic clusters

#### "normal" matter clumps



#### dark matter clumps

### Visual Evidence of Park Matter

\* The hot gas in each cluster was slowed by a drag force, similar to air resistance, during the collision. In contrast, the dark matter was not slowed by the impact because it does not interact directly with itself or the gas except through gravity. Therefore, during the collision the dark matter clumps from the two clusters moved ahead of the hot gas, producing the separation of the dark and normal matter seen in the image

All three methods of measuring cluster mass indicate similar amounts of dark matter: There is about 5 to 6 times more dark matter than luminous one



### Dark Matter Filaments

- More recent observations shows that clusters of galaxies are connected with dark matter filaments
- \* Theoreticians had predicted this: they stated that
- dark matter is distributed in our Universe along a gigantic web of interconnected filaments (which also includes normal ionized matter emitting Xrays.) The filaments' mass is at least 90% dark matter
### Dark matter filament linking two galactic clusters



### Dark matter web linking clusters of galaxies



#### Theoretical (computed) image

### What might dark matter be made of?

- \* We cannot expect to be able to visually see all ordinary matter, some is too faint to be detected
- \* Park matter, then, can be split in two
  - 1. Ordinary dark matter
  - 2. Extraordinary dark matter

# Ordinary Matter

 Protons and neutrons (which make up most of the mass of ordinary matter) belong to a category of particles called baryons (made of 3 quarks)

\* Ordinary matter is also called baryonic matter (fermions include baryons and leptons [electrons and neutrinos])

# Extraordinary Matter



### \* Neutrinos and free electrons

### \* Unknown:

\* Extraordinary dark matter and black holes (we do not know what is inside a black hole - maybe quark soup?)

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### Park Matter Two Basic Options

# 1. Ordinary Park Matter, made of MACHOs

# 2. Extraordinary Park Matter, made of WIMPs

# Ordinary Park Matter

- \* MACHOs: Massive Compact Halo Objects
  - trillions of dead or failed stars in the halos of galaxies
  - 2. trillions or more planets
- \* all are too faint to be detected with our current instruments, though we are getting better at detecting them "nearby"



MACHOs occasionally make other stars appear brighter through lensing

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MACHOs occasionally make other stars appear brighter through lensing

... but not enough lensing events to explain dark matter

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# Extraordinary Park Matter

- \* WIMPS: Weakly Interacting Massive Particles
  - \* mysterious neutrino-like particles but much heavier
  - \* sometimes called cold dark matter to separate them from the faster moving neutrinos

# Why Believe in WIMPs?

1. There's not enough ordinary matter to explain why galaxies keep whole

2. Models involving WIMPs explain how galaxy formation works

3. WIMPs could be leftovers from Big Bang

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# Snapshot



### \* What is the evidence for dark matter in galaxies?

\* The orbital velocities of stars and gas clouds in galaxies do not change much with distance from the center of the galaxy. Applying Newton's laws of gravitation and motion to these orbits leads to the conclusion that the total mass of a galaxy is far larger than the mass of its stars. Because no detectable visible light is coming from this matter, we call it dark matter





\* What is the evidence for dark matter in clusters of galaxies?

\* We have three different ways of measuring the amount of dark matter in clusters of galaxies: from galaxy orbits, from the temperature of the hot gas in clusters, and from the gravitational lensing predicted by Einstein. All of these methods agree that the total mass of a cluster is about 50 times the mass of its stars, implying huge amounts of dark matter





- \* What is the evidence for dark matter filaments linking clusters of galaxies?
- \* Gravitational lensing and a spike in X-ray emissions along filaments (due to an excess of ionized ordinary matter being pulled by gravity toward the filaments)
- \* All estimations calculate that 90% or more of the filaments' mass is dark matter



\* Does dark matter really exist?

\* Yes. We now have direct evidence of dark matter seen in the collision of two galaxy clusters. The dark matter is "seen" via the gravitational lensing it creates on objects far behind it and in our line of sight

## Snapshot



\* What might dark matter be made of?

Some of the dark matter could be ordinary or baryonic matter in the form of dim stars or planetlike objects, but there does not appear to be enough ordinary matter to account for all the dark matter. Most of it is probably extraordinary or nonbaryonic matter consisting of undiscovered particles that we call WIMPs

# The role of dark matter in galaxy formation

- Park matter's gravity was probably the main force that caused protogalactic clouds to become galaxies and galaxies to group into clusters of galaxies as
- \* the hydrogen and helium gas in the protogalactic clouds collapsed inward and gave birth to stars and

# \* dark matter couldn't radiate its energy and so remained distributed throughout the cluster

### Gravity of dark matter is what caused protogalactic clouds to contract early in time





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Dark matter is still pulling things together

After correcting for Hubble's Law, we can see that galaxies are flowing toward the densest regions of

space

# The largest known structures in the Universe

 Beyond 300 million light-years from us, the gravitational tugs of galaxies becomes insignificant

 Hubble's law becomes the primary method of determining distances to galaxies

# Superclusters and Supervoids

- \* We saw how galaxies come in groups and clusters
- \* Clusters like company: they form superclusters (clusters of clusters)
- \* An "average" supercluster has upward of 50,000 galaxies contained within a diameter of some 200 to 300 million light-years

# Superclusters and Supervoids...

- Superclusters are not evenly distributed in space: they form supercluster complexes which are separated by huge voids, called supervoids
- \* Yet superclusters are linked (filament-like) across the sky on the "surface" of the supervoids
- \$ 90% of the Universe appears to consists of voids while the remaining 10% seems to be superclusters linked by a web of filaments

### A 3-D map of the nearest Superclusters

**Clusters and** superclusters Coma Abell 1367 nearby Abell 779 75 224 Abell 569 Hydra Abell 3581 Virgo Camelopardalis Perseus Local Centaurus Group Fornax-. Eridanus Pavo 75 206 Pegasus Abell 194 5,000 Abell 2911 1.000 5550 Abell 2870 899 DATA ON **Our Local** Group is gravita-1.aan tionally pulled toward the Virgo cluster, which is part of the Norma-Hydra-600 6.000 Centaurus Supercluster. Astronomy: Roen Kelly; after M. Hudson

### A map of the nearest Superclusters and Supervoids



#### Maps of galaxy positions reveal extremely large structures: superclusters and supervoids



#### Notice how the structure is organized close to the present and gets more and more uniform as we look to the past





Size of expanding box in millions of light-years

Models show that gravity of dark matter pulls mass into denser regions (the Universe grows lumpier with time)

Dark matter creates a Universe-sized scaffolding (or web)

#### Galaxies appear to be arranged in immense structures hundreds of millions of light-years across



# The Shape of Park Matter's Distribution

- \* Using gravitational lensing effects, researchers made detailed measurements of the spatial distribution of dark matter in about 20 massive galactic clusters
- \* The team obtained clear evidence that the distribution of dark matter in the clusters has, on average, an extremely flattened shape rather than a simple spherical contour (or flat string-like)

# The Shape of Park Matter's Distribution...



# The Shape of Park Matter's Distribution...

\* This research strongly supports the prevailing dark matter model, which begins with the assumption that dark matter consists of weakly interacting massive particles that are relics of the Big Bang

\* These particles are assumed to be "cold" they have negligibly small thermal motions



\* What is the role of dark matter in galaxy formation?

\* Because most of a galaxy's mass is in the form of dark matter, the gravity of that dark matter is probably what formed protogalactic clouds and then galaxies from slight density enhancements in the early Universe



### \* What is origin of dark matter?

\* The prevailing theory is that dark matter consists of weakly interacting cold massive particles that are relics of the Big Bang



- \* What are the largest structures in the Universe?
- Galaxies appear to be distributed in gigantic chains and sheets that surround great voids, or supervoids. These giant structures trace their origin directly back to regions of slightly enhanced density early in time

# The Universe's Fate

\* With the addition of dark matter which adds to the force of gravity and dark energy which fights against it, what is the future like?

\* Will the Universe continue expanding forever? Or will it collapse?

\* Is the expansion of the Universe accelerating? What is gravity doing?
### Will the Universe continue expanding forever?

#### \* How will the Universe end?

- a. With enough gravity, the expansion will someday halt and the Universe will start a collapse (ending in a Big Crunch!)
- b. Else, the Universe will keep on expanding forever, get colder and matter will eventually "evaporate" as the porosity of space will not allow particles to exist





Poes the Universe have enough kinetic energy to escape its own gravitational pull?

When the Universe's density is larger than the critical density... it all ends in a Big Crunch!

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Poes the Universe have enough kinetic energy to escape its own gravitational pull?

When the Universe's density is less than the critical density... it expands forever

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Fate of Universe depends on amount of dark matter The impact of gravitational force depends on the density of matter in the Universe coasting Universe Critical density: 8x10-30 gram per cm<sup>3</sup> **Big Bang** not enough

dark matter

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#### Fate of Universe depends on amount of dark matter

Assuming no dark energy presence, luminous matter is 0.5% of the total energy density needed to halt the expansion. Is there 200 times more dark matter than luminous one? We only find about 55 times. So we would assume that the Universe will expand forever



not enough dark matter Amount of dark matter is ~25% of the critical density suggesting fate is eternal expansion

but...

# The expansion of the Universe is accelerating

- Observations of distant white dwarf supernovae seem to indicate the Universe's expansion is not coasting but accelerating
- \* A mysterious repulsive force is needed to explain this acceleration: dark energy

\* This repulsive force only acts at galacticlike distances



- \* Imagine throwing a rock upward from where you stand
- Depending on its initial velocity, it will go up a certain distance, slow down and then come back down
- \* What would you think if once you have thrown the rock upward, its velocity would not slow down but get faster instead? What would cause that?

#### Using white-dwarf supernovae as standard candles



From the white-dwarf supernova's measured brightness, we find the distance to a very distant galaxy which is usually too faint for us to get a good measurement on its distance

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#### Brightness of distant white-dwarf supernovae tells us how much Universe has expanded since they exploded

Distant galaxies before supernova explosions



#### The same galaxies after supernova explosions



### Distances from Hubble Law and White Dwarf Supernovae

Comparing the distances obtained via calculating the absolute magnitudes of these supernovae and the distances obtained from their galaxies red-shifts we find differences that increase with distance

\* This indicates that Hubble's Law is not linear: the expansion's speed is not constant

#### Accelerating Universe is best fit to supernova data



Universe expansion appears to be speeding up!

**Observations** of distant white dwarf supernovae seem to indicate the Universe's expansion is not coasting but accelerating

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#### The Fate of the Universe depends on amounts of dark matter and dark energy



#### The more energy there is in the Universe, the older it is



### The Age of the Universe depends on amounts of dark matter and dark energy



# Age of the Universe

#### Latest studies (Mar. 2013) coming out of detailed analysis of the Wilkinson Microwave Anisotropy Probe indicate that the Universe is 13.798 billion years old

\* The error range is +/- 37 million years



- Suppose that the Universe has more dark matter than we think there is today - how would that change the age we estimate from the expansion rate
  - A. Estimated age would be larger (older)
  - B. Estimated age would be the same
  - C. Estimated age would be smaller (younger)

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#### Dark Energy is not a "recent" force

#### Host Galaxies of Distant Supernovae

HST - ACS/WFC



NASA, ESA, and A. Riess (STScI)

STScI-PRC06-52

#### These 5 supernovae are between 3.5 to 10 billion lightyears away Measurements indicate that the dark energy was present 10 billion years ago. It is not a new constituent of space



Without dark energy being present, the mass density set here is a closed Universe

However, with a minimal dark energy amount, the Universe is accelerating

Hence if dark energy exists, even a little, the Universe expands forever

### Evidence for Vark Energy

- 1. Type la supernovae
- 2. Presence detected as far as 10 billion years ago
  - And, to be seen in the next lecture:
- 3. Cosmic Microwave Background (CMB, WMAP)
- 4. Gravitational Lensing
- 5. Large scale structure of the Universe

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### What is "Dark Energy"?

- \* Empty space has a fundamental energy: empty space is not "nothing" (Einstein)
- \* It is not dense: 10<sup>-29</sup> g/cm<sup>3</sup>
- \* Its strength gets effective only at great cosmic distances (sum effect)
- \* It is also called the vacuum energy
- It applies a negative pressure on space (positive pressure: compression; negative: repulsion = expansion)

## What is "Dark Energy"?

- \* Two more possibilities (but both raise more questions)
- 1: QM states that empty space is full of virtual particles that continually appear and disappear - but energy calculations are off by 10<sup>120</sup> times too large!
- 2: Park energy is a kind of dynamical negative energy field which fills space (called "Quintessence") and is similar in concept to the Higgs field. No one knows why it would exist

# What is "Park Energy"?

- \* Could Einstein Theory of Gravity be wrong?
- \* That would not only affect the expansion of the Universe, but it would also affect the way that normal matter in galaxies and clusters of galaxies behaves
- \* There are candidate theories, but none are compelling
- \* The research for Dark Energy continues!





#### \* Will the Universe continue expanding forever?

\* Even before we consider the possibility of a mysterious dark energy, the evidence points to eternal expansion. The critical density is the average matter density the Universe would need for the strength of gravity to eventually halt the expansion. The overall matter density of the Universe appears to be only about 25% of the critical density needed to stop the expansion

### Snapshot

### \* Is the expansion of the Universe accelerating?

galaxies

between

distance

supernova data

10

0

past <del>< \_;</del>→ future

time (billions of years from now)

-4.4

\* Observations of distant supernovae indicate that the expansion of the Universe is speeding up. No one knows the nature of the mysterious force (powered by the dark energy) that could be causing this acceleration



- \* The Cosmic Web
  - http://www.youtube.com/watch?v=74lsySs3RGU
- \* Large Scale Structure of the Universe
  - http://www.youtube.com/watch?v=VJpC\_oQQxPI
- \* 3-D simulation large scale structure of the Cosmos
  - http://www.youtube.com/watch?v=FFlzyxSQhTc
- \* How do we know the Universe is flat?
  - http://www.youtube.com/watch?v=zqb1lSdqRZY